

# Access to European High Performance Computing – A PRACE User Perspective

Åke Nordlund

Niels Bohr Institute  
University of Copenhagen

# Overview

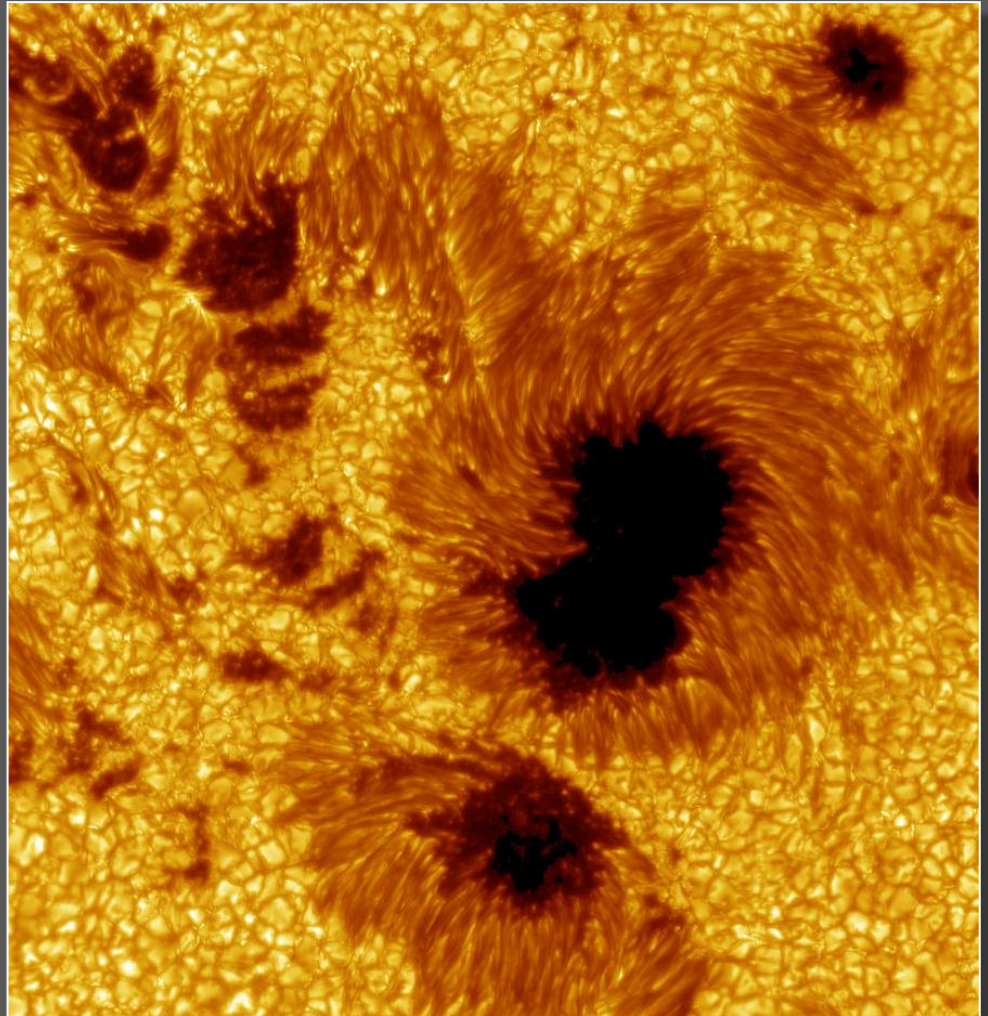
- ▣ Context
  - User profile – experiences
- ▣ PRACE Projects
  - Solar research
  - Other grant examples: The supernova project
- ▣ PRACE-related e-Infrastructure Requirements
  - Tier-0 / Tier-1 access
  - Data *analysis*, *archival*, and peer *access*

# User Profile

- ▣ 40 years HPC experience
  - 1970's: UNIVAC, CDC
  - 1980's: IBM, Convex, ...
  - 1990's: Connection Machine (CM-200)
  - 2000's: NASA/Ames, Stuttgart, Jülich (JUGENE)
  
- ▣ Danish Center for Scientific Computing
  - DCSC-KU: Intel-cluster, GPU-cluster (200 GPUs)
  - ScaleMP Virtual Shared Memory system
    - ▣ 'Analysis server'

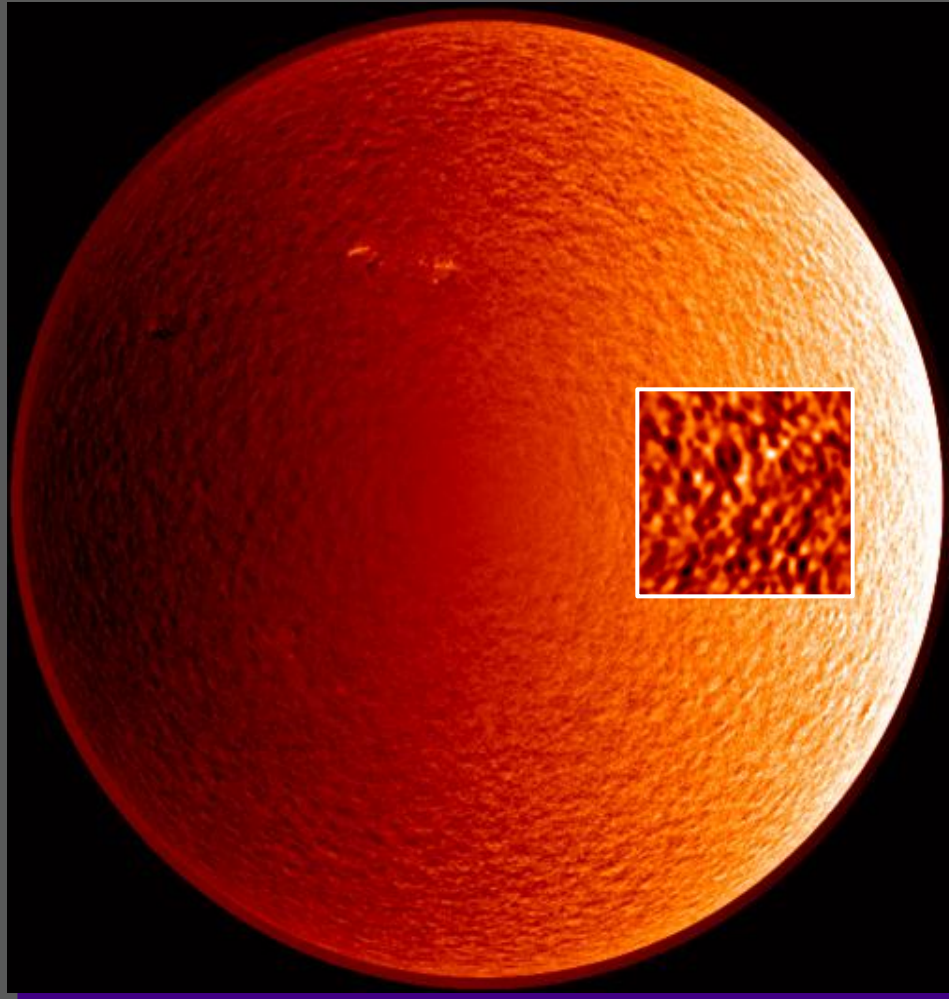
# Solar Modeling

The magnetic field in the solar corona is controlled by the magnetic field at the surface of the Sun



# A Nearly Scale Free Spectrum!

Doppler Image of the Sun

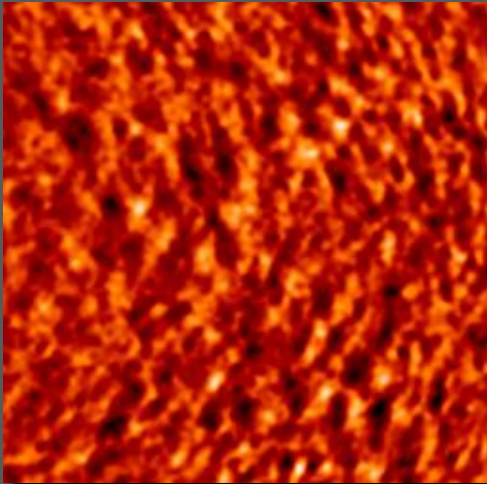




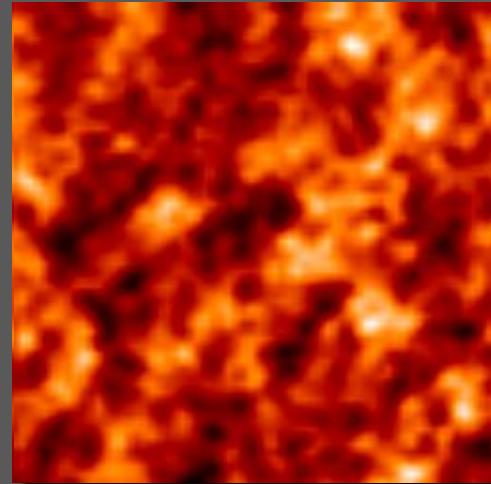
# Solar horizontal velocity (observed)

Scales differ by factor 2 – which is which?

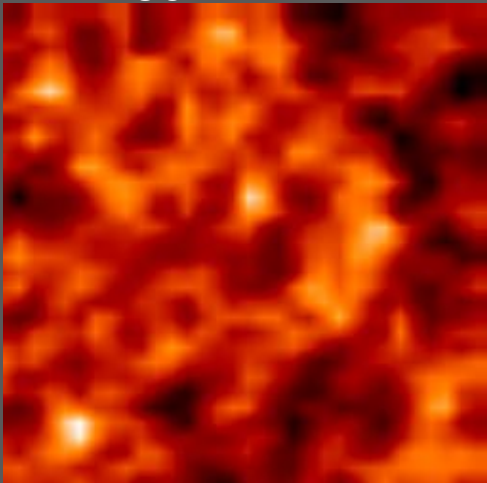
400 Mm



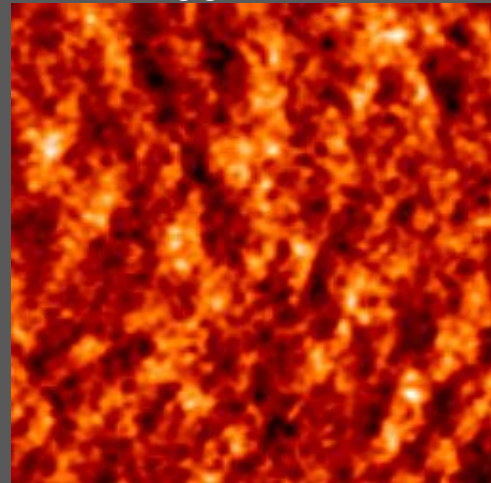
100 Mm



50 Mm



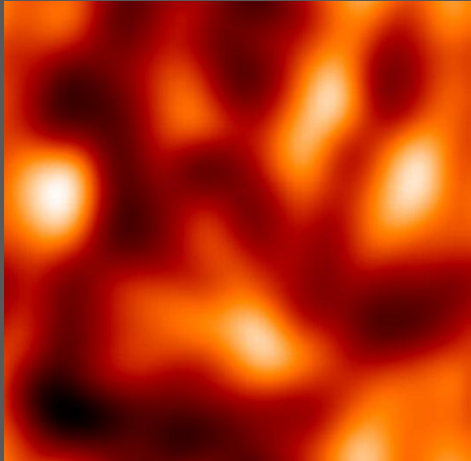
200 Mm



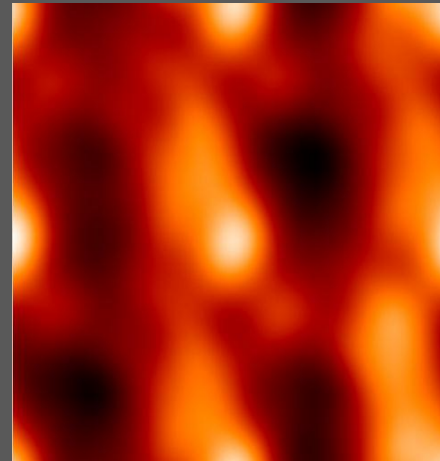
# Solar horizontal velocity (model)

Scales differ by factor 2 – which is which?

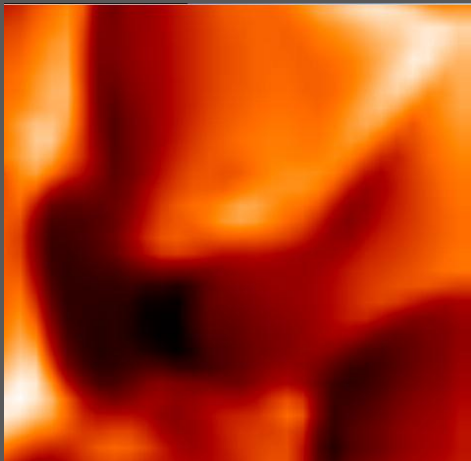
12 Mm



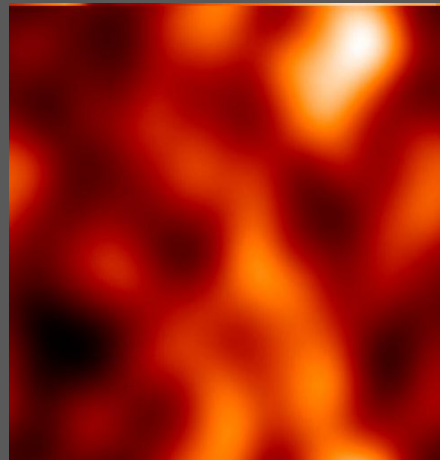
24 Mm



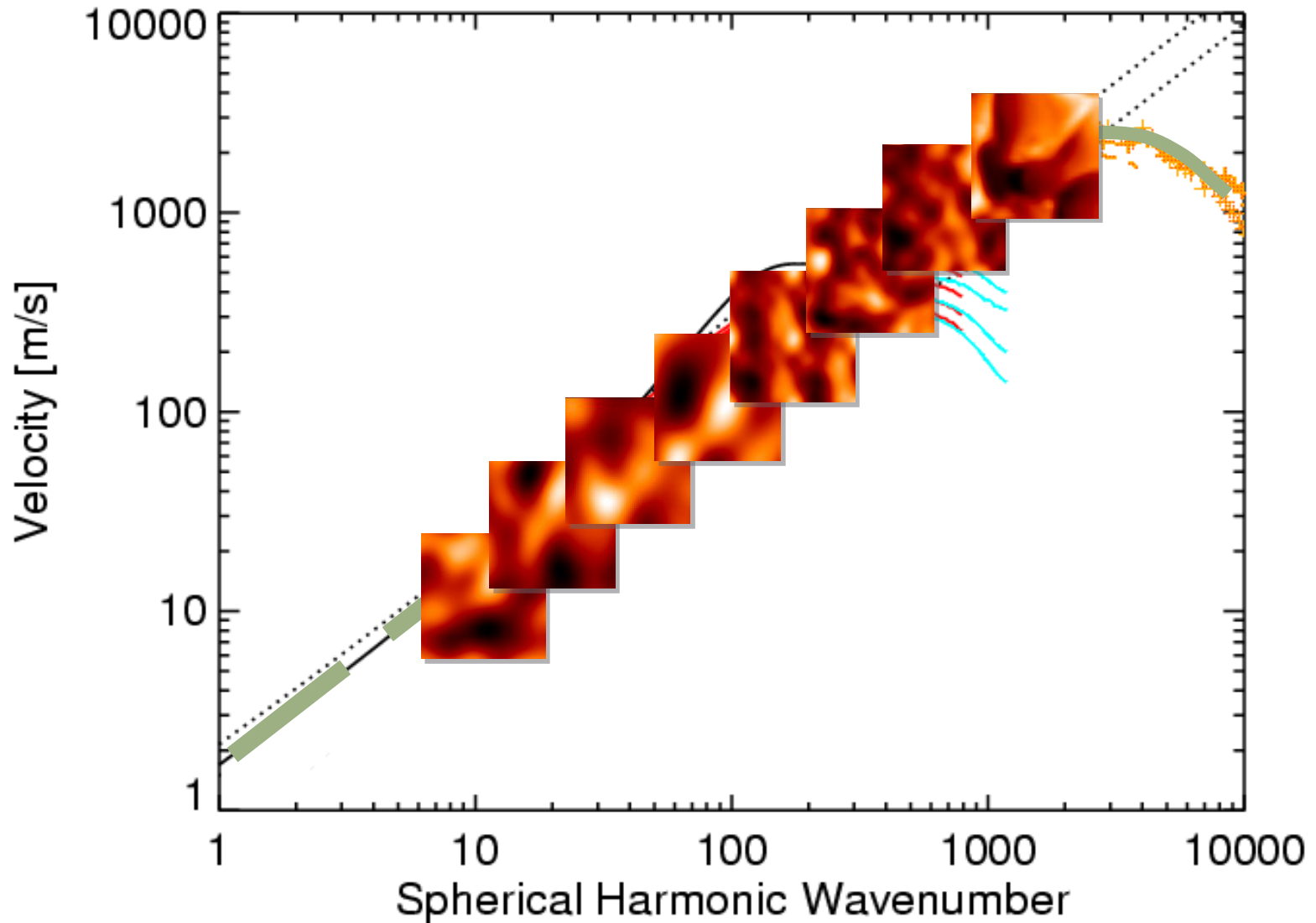
3 Mm



6 Mm



# Solar velocity spectrum

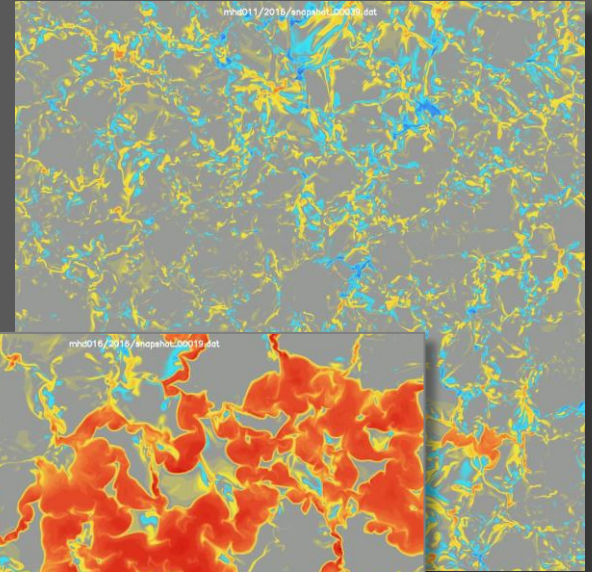




# Solar Active Region 3-D Simulations

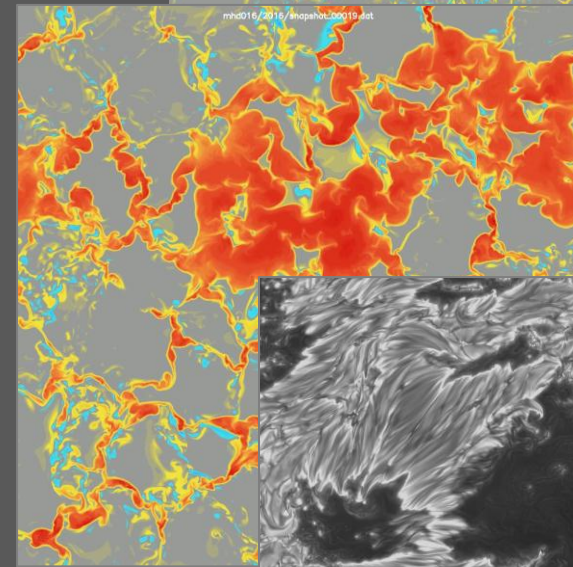
## ▣ 'Quiet' Sun

- Zero mean field with  $\langle B^2 \rangle^{1/2} \sim 50 - 150 \text{ G}$



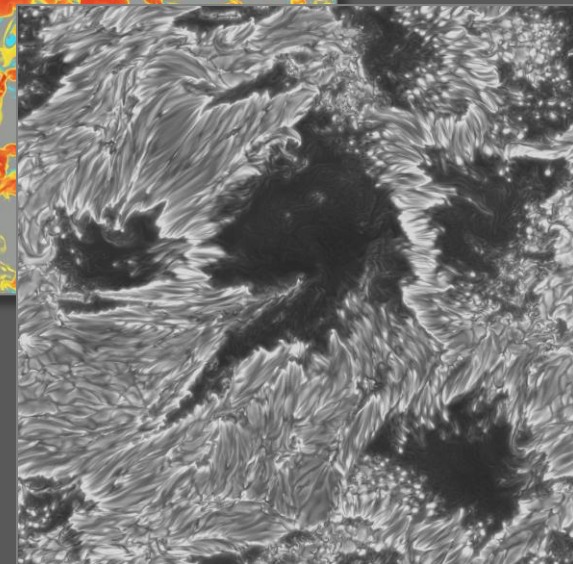
## ▣ Plage Region

- Mean vertical field  $B \sim 600 \text{ G}$
- 24,000 km size  
grid sizes **down to 6 km**



## ▣ Active Region with Sunspots

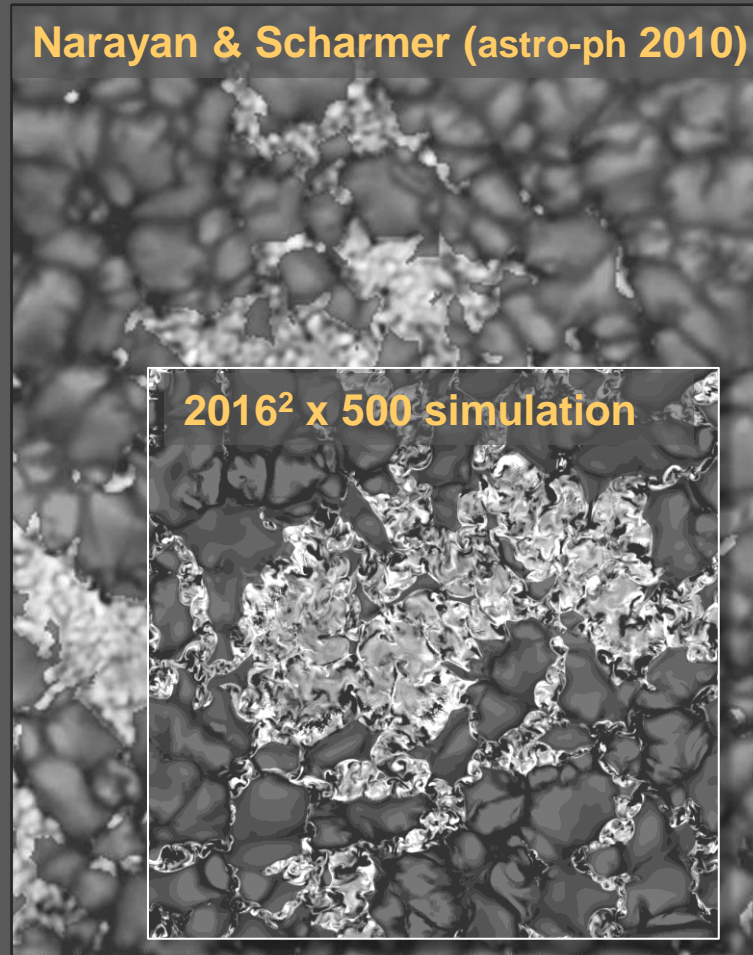
- Zero mean field with  $\langle B^2 \rangle^{1/2} \sim 1.5 \text{ kG}$
- 48,000 km size,  
horizontal grid size 24 km



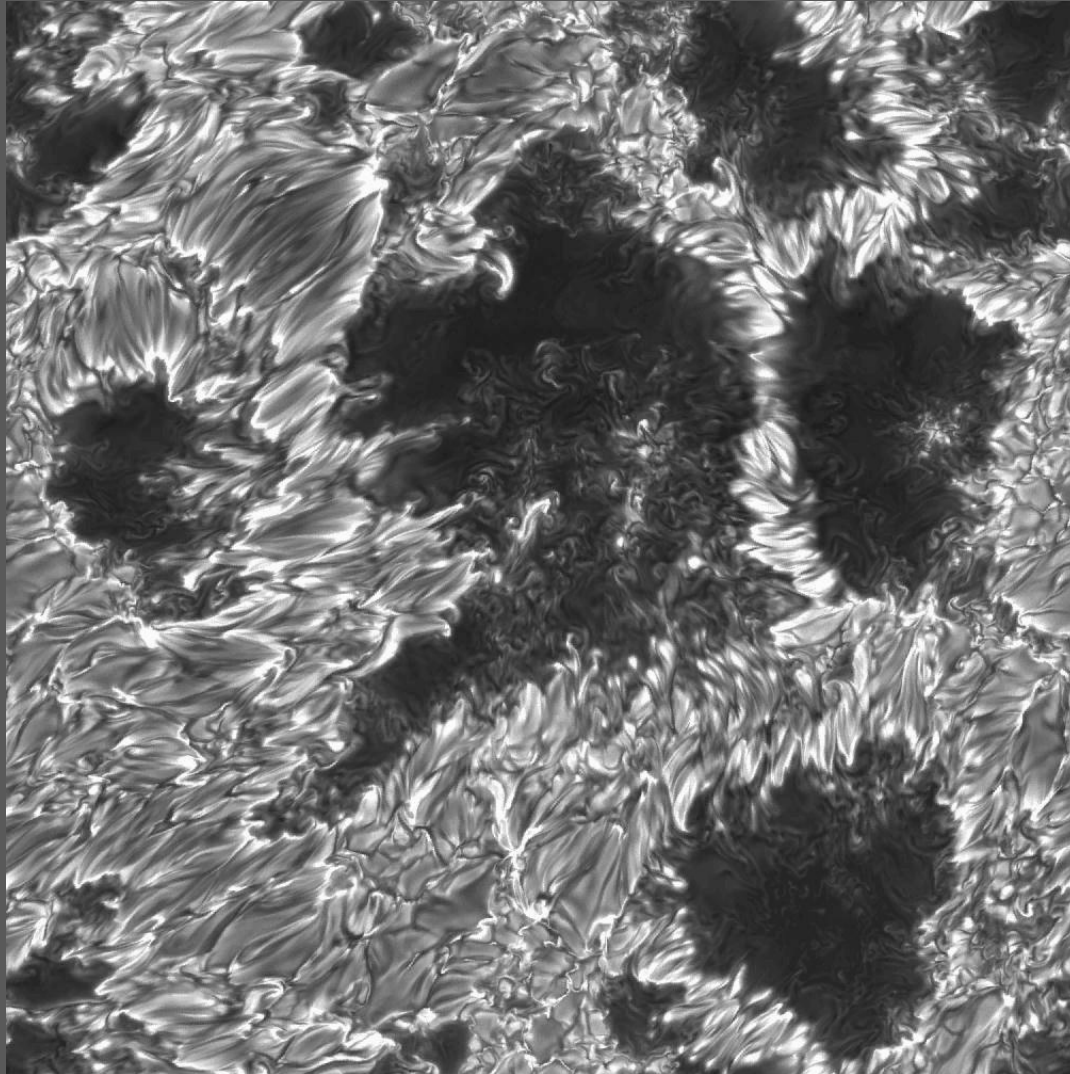
# Solar 'Plage' Regions

Line-of-sight velocity

Narayan & Scharmer (astro-ph 2010)



# Continued spot evolution (simulations at NASA/Ames by Bob Stein)

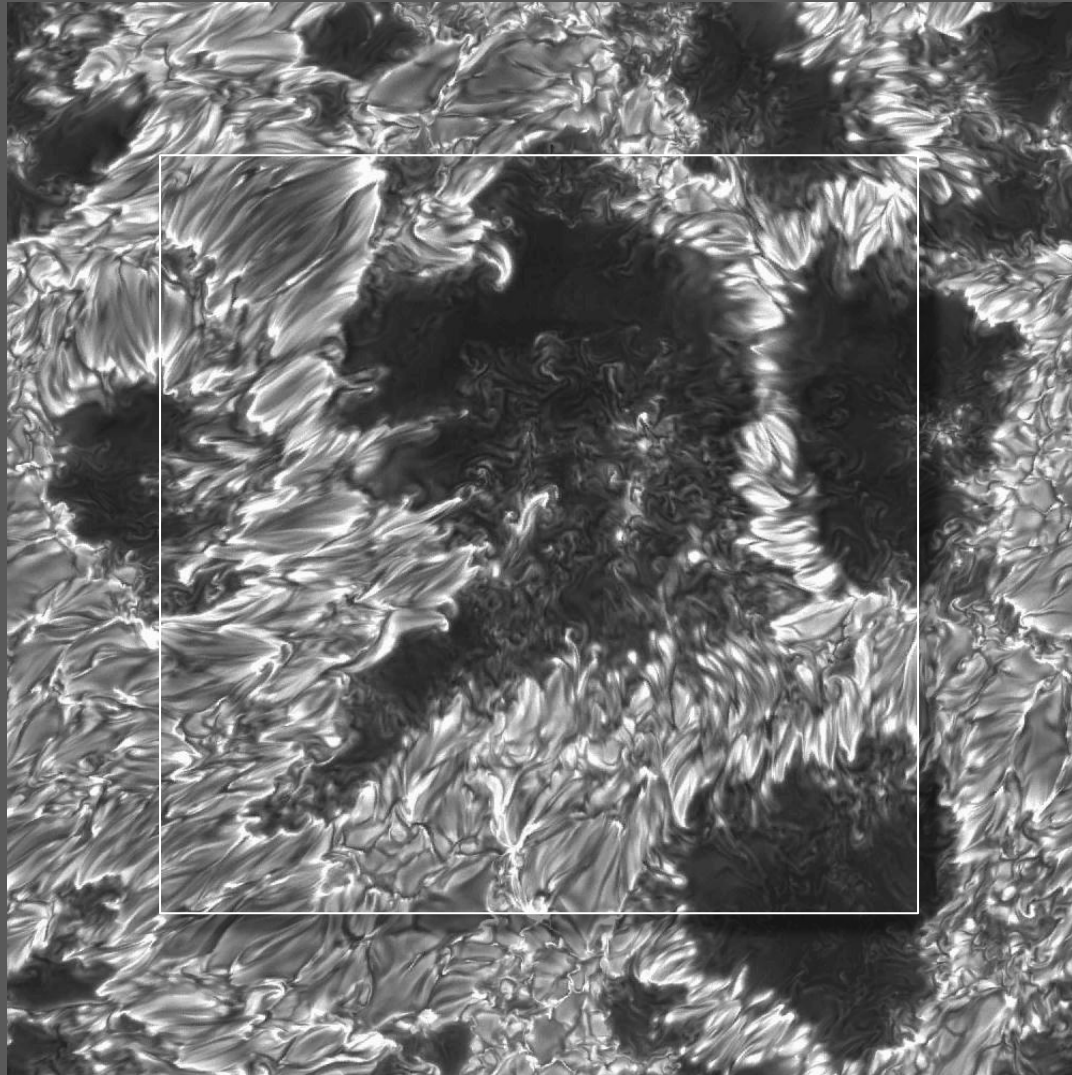


Size:  $48^2 \times 10$  Mm  
Mesh:  $2016^2 \times 500$



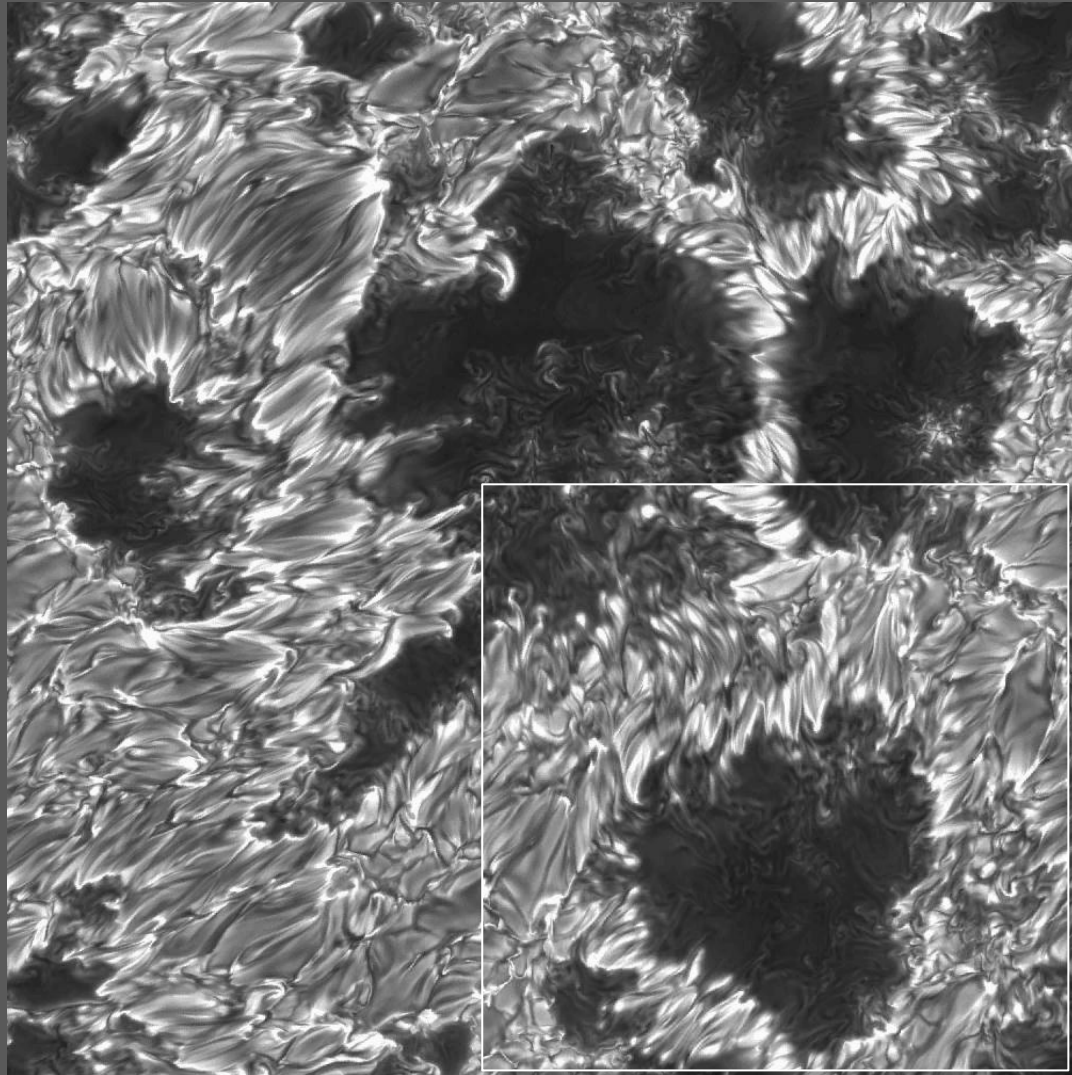
# Continued Spot Evolution

## Zoom in on the central spot



# Continued Spot Evolution

## Zoom in on the spot at lower right





# Importance and Impact of Visualization

## ▣ Importance:

- Visualization in 3-D, including time animation is crucial
- Complements quantitative / statistical analysis

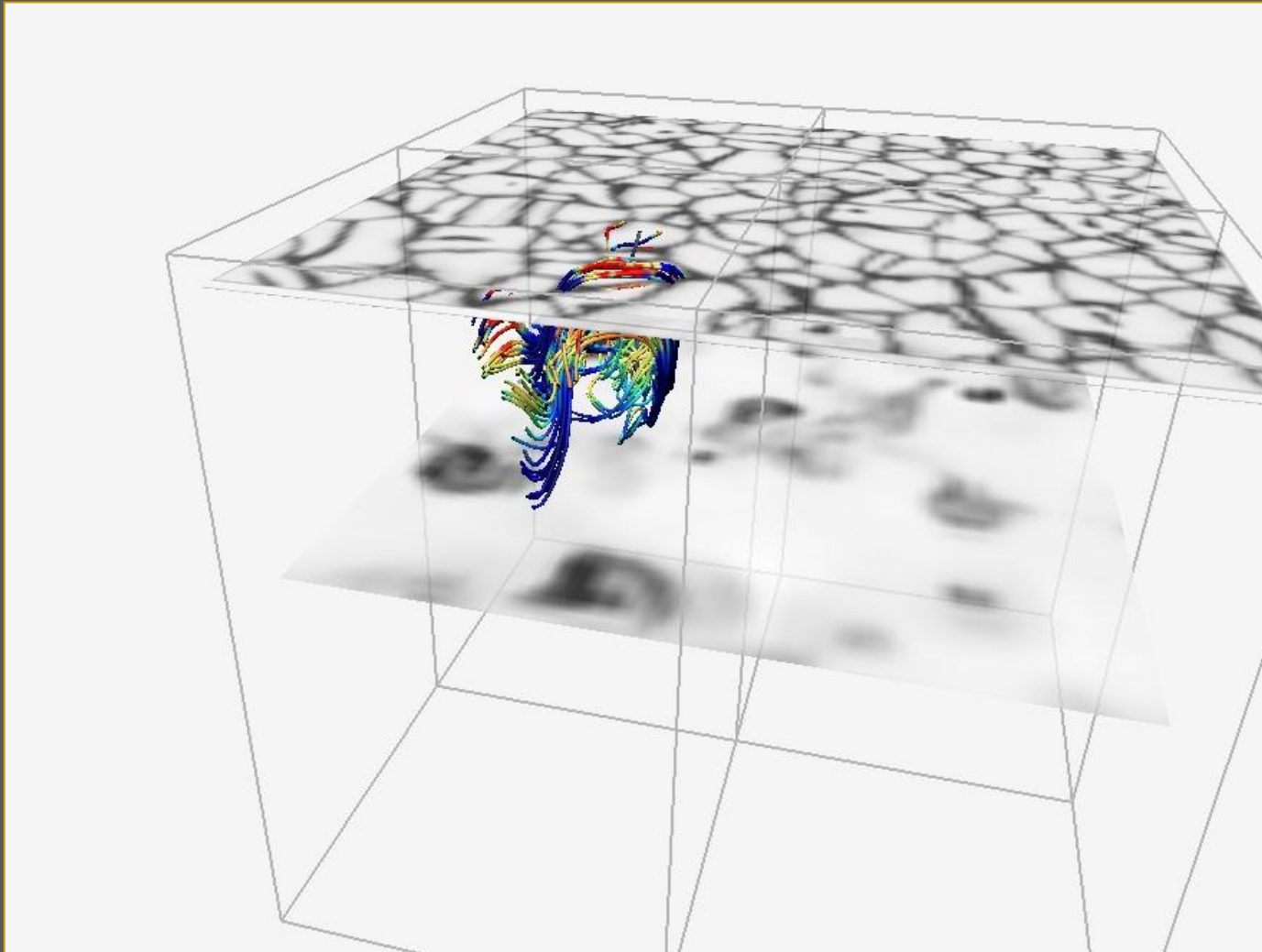
## ▣ Impact:

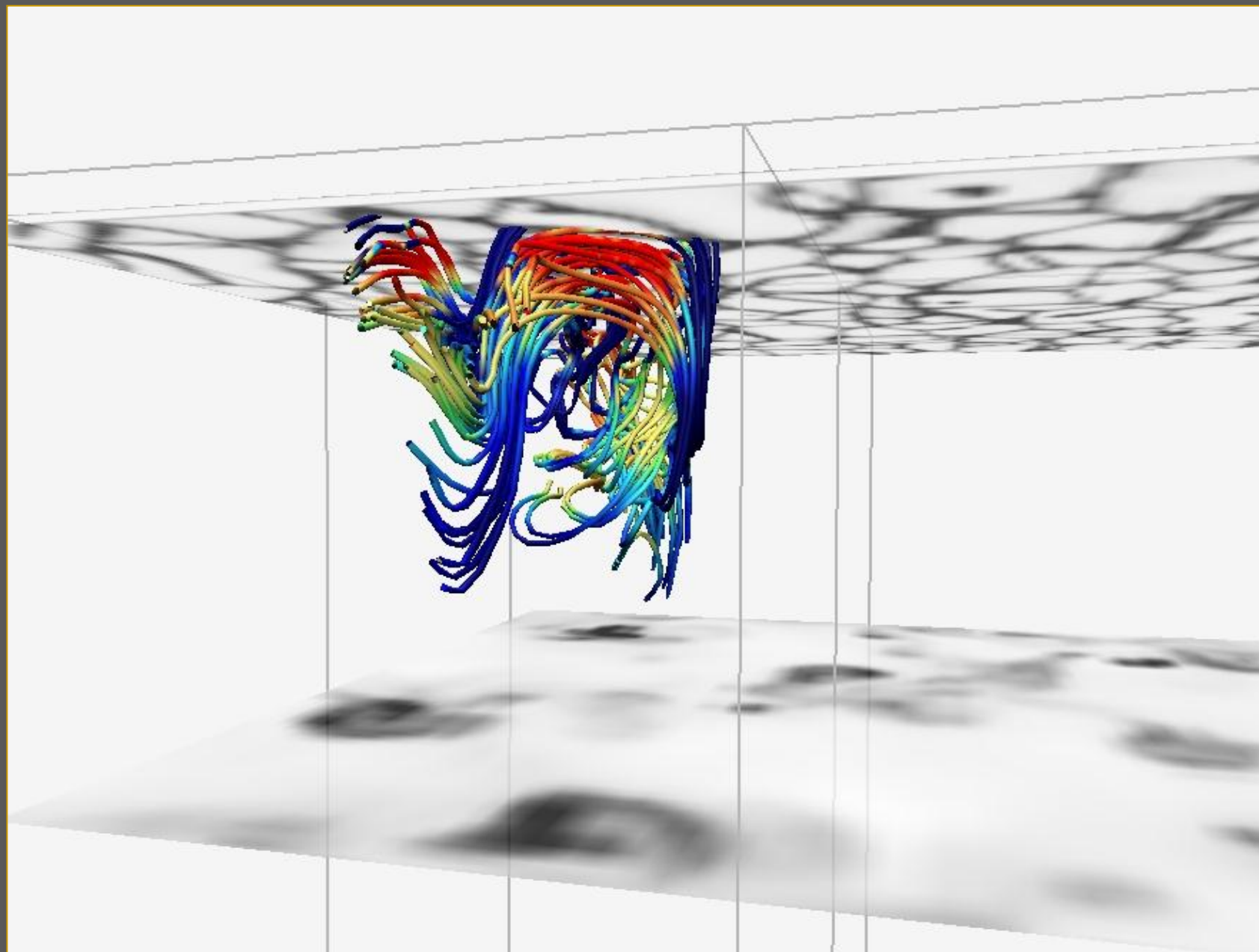
- High cadence fluid snapshots, each  $\sim 0.1\text{-}0.2$  TB
- High cadence sub-sampled particle snapshots, each  $\sim 1$  TB
  - ▣ Restart snapshots: 30 TB

## ▣ Requirements:

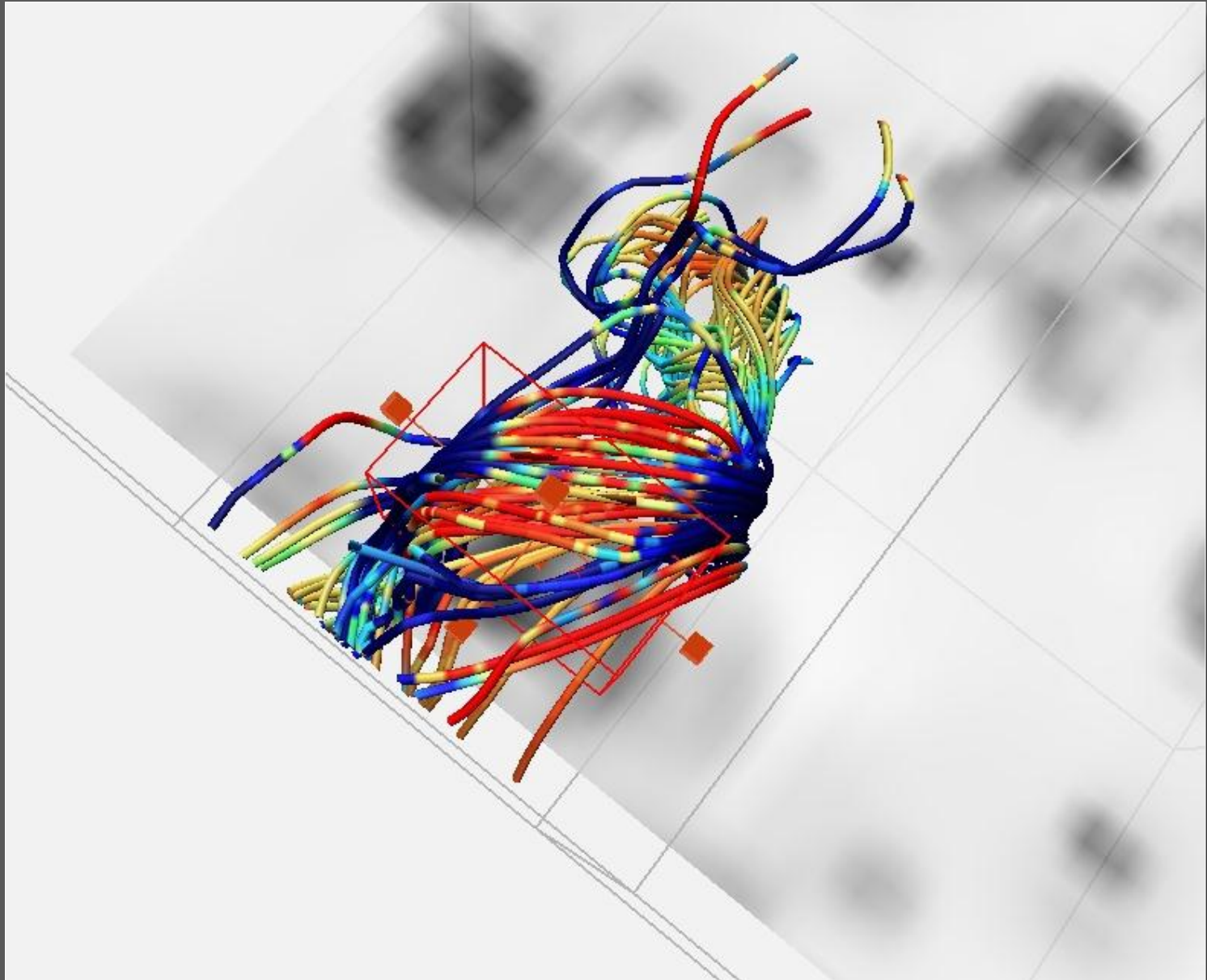
- High data I/O bandwidth!!
- Wavelet compression from server to user

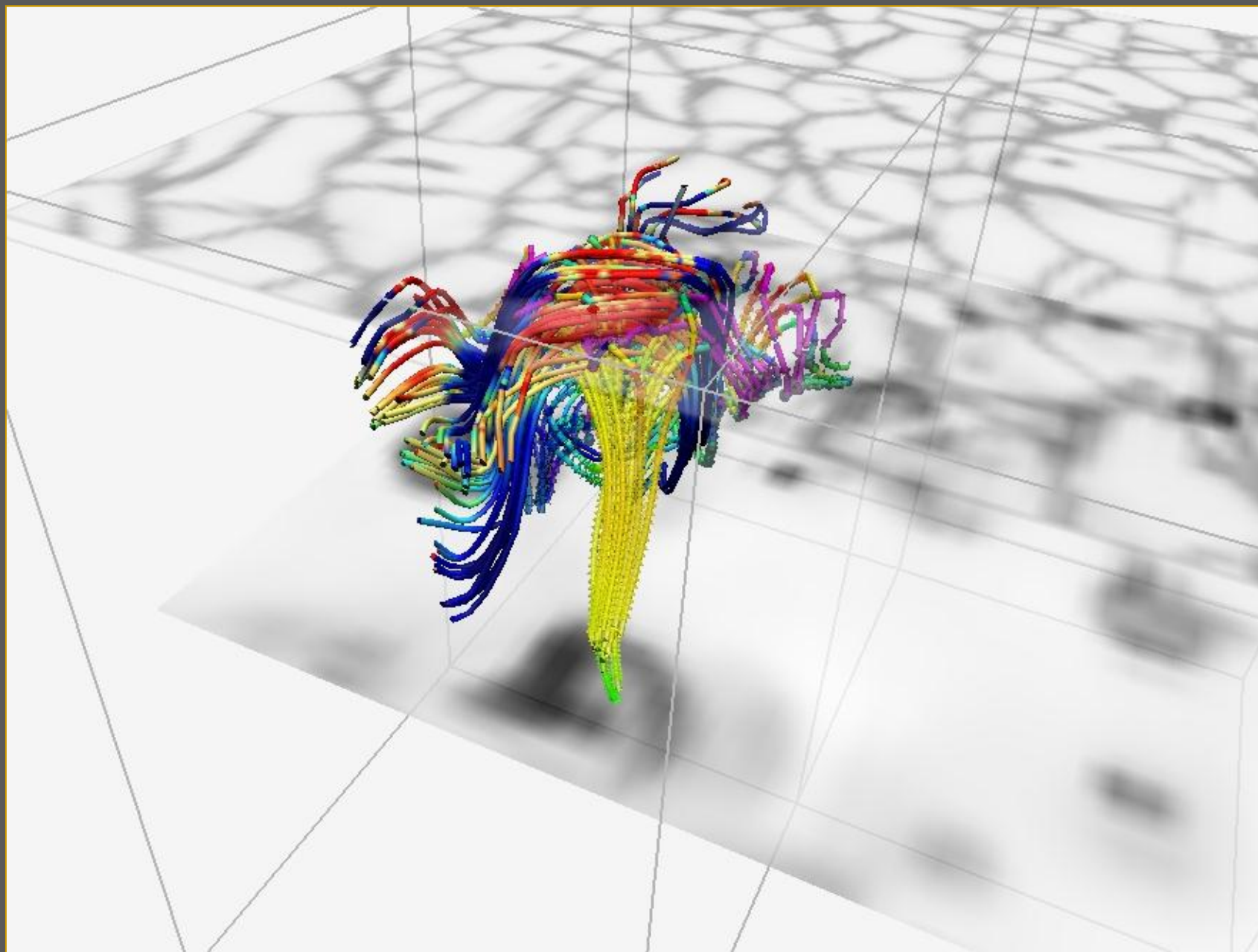
# 3-D (NCAR/Vapor) visualizations, illustrating the process



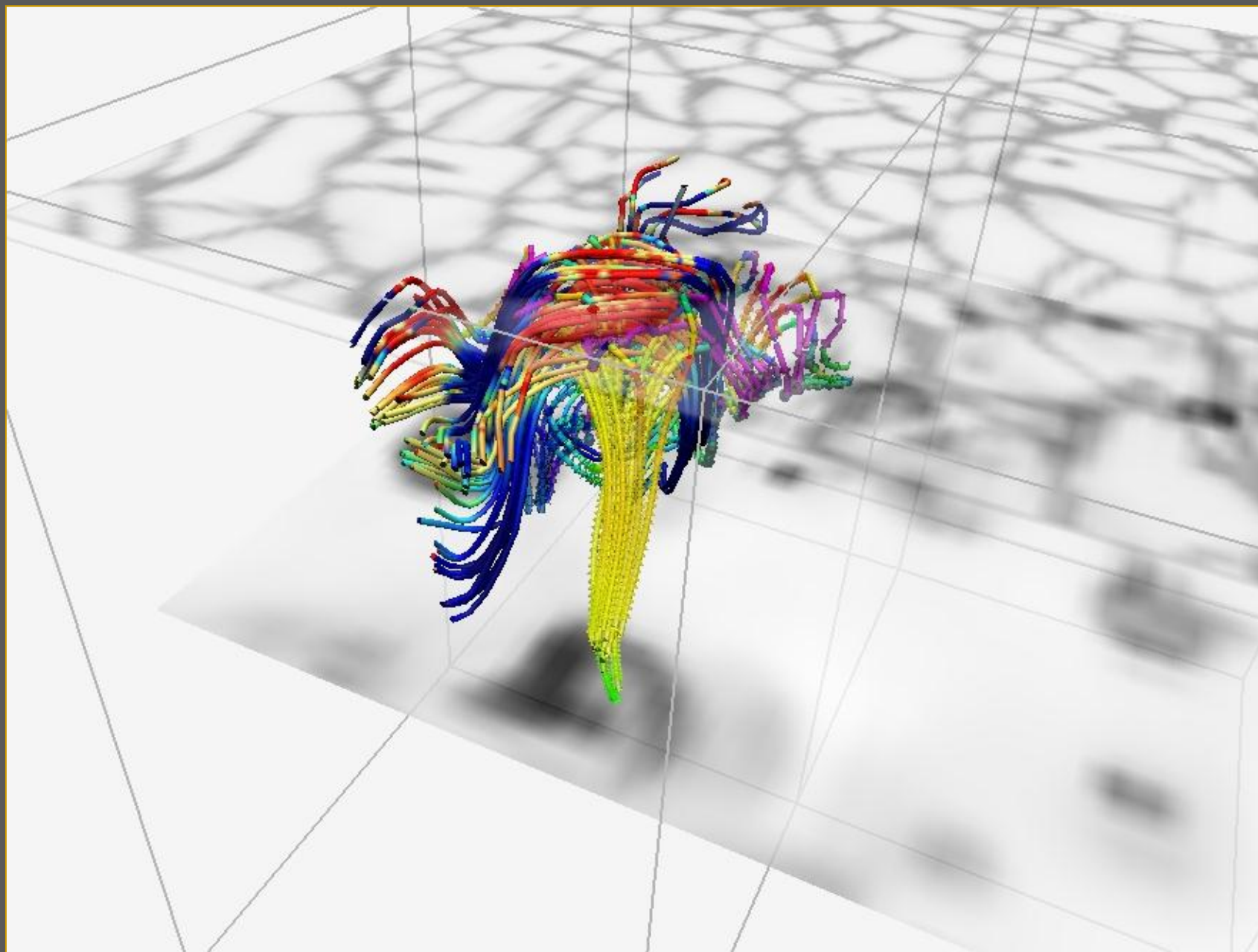


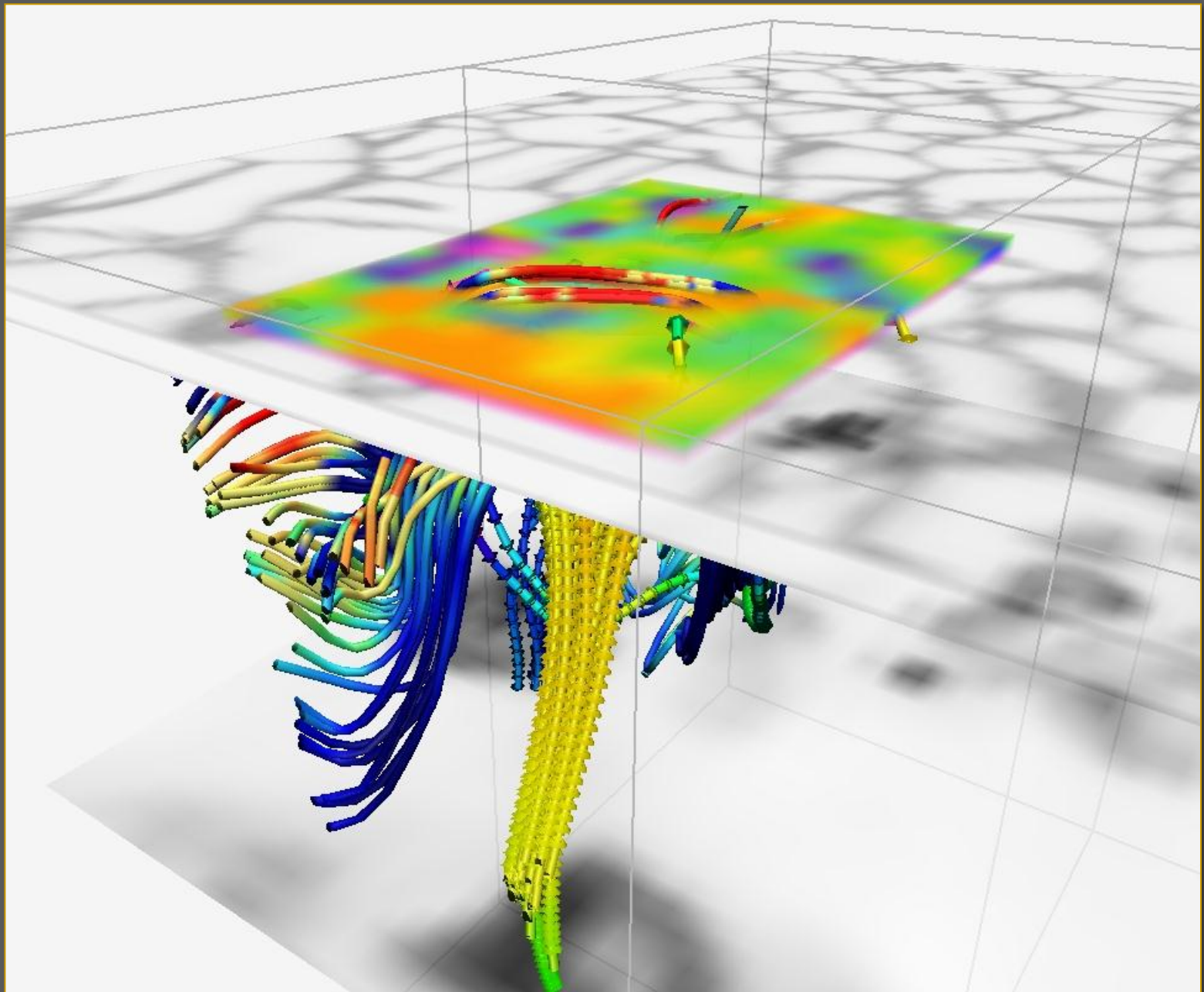
# View from above

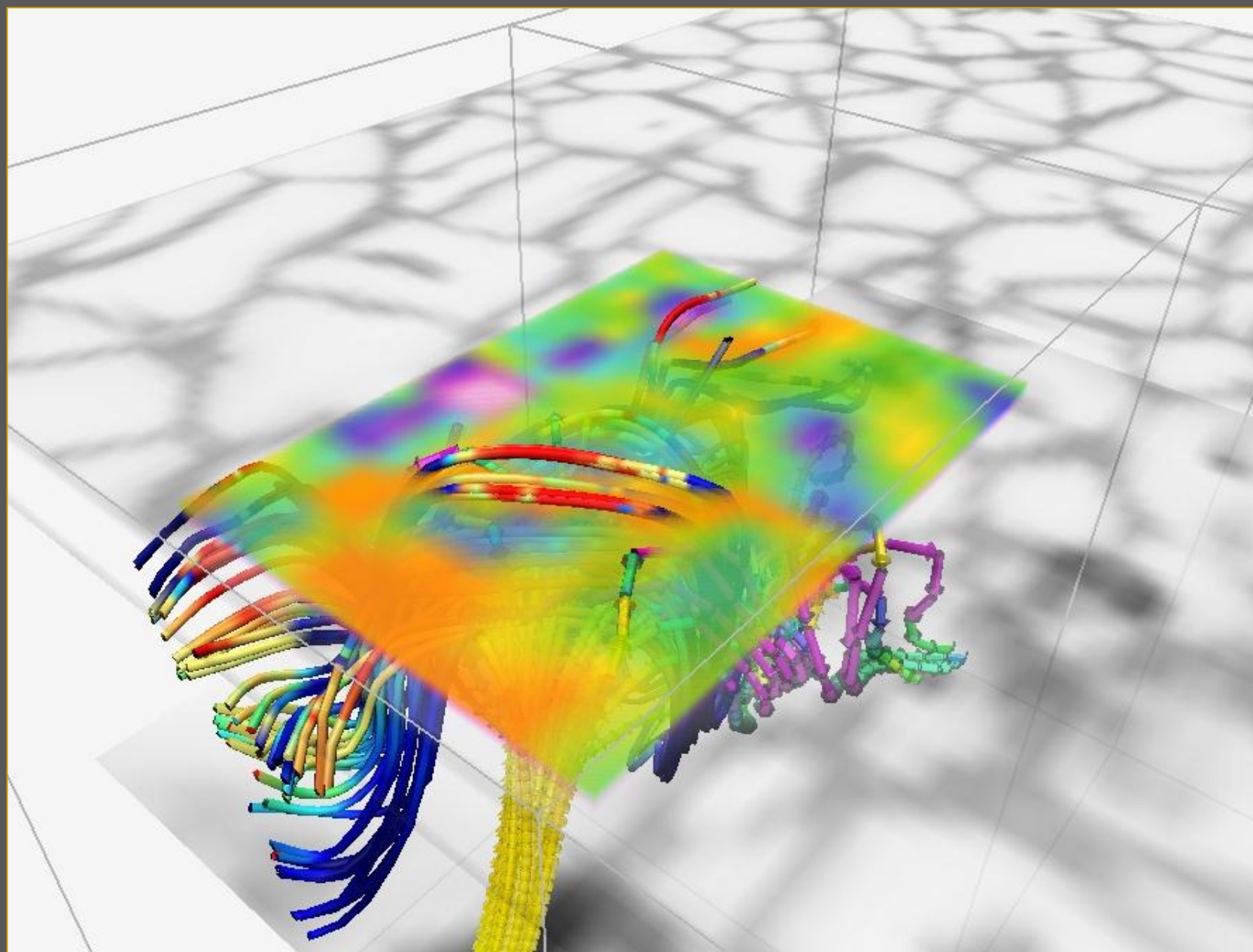








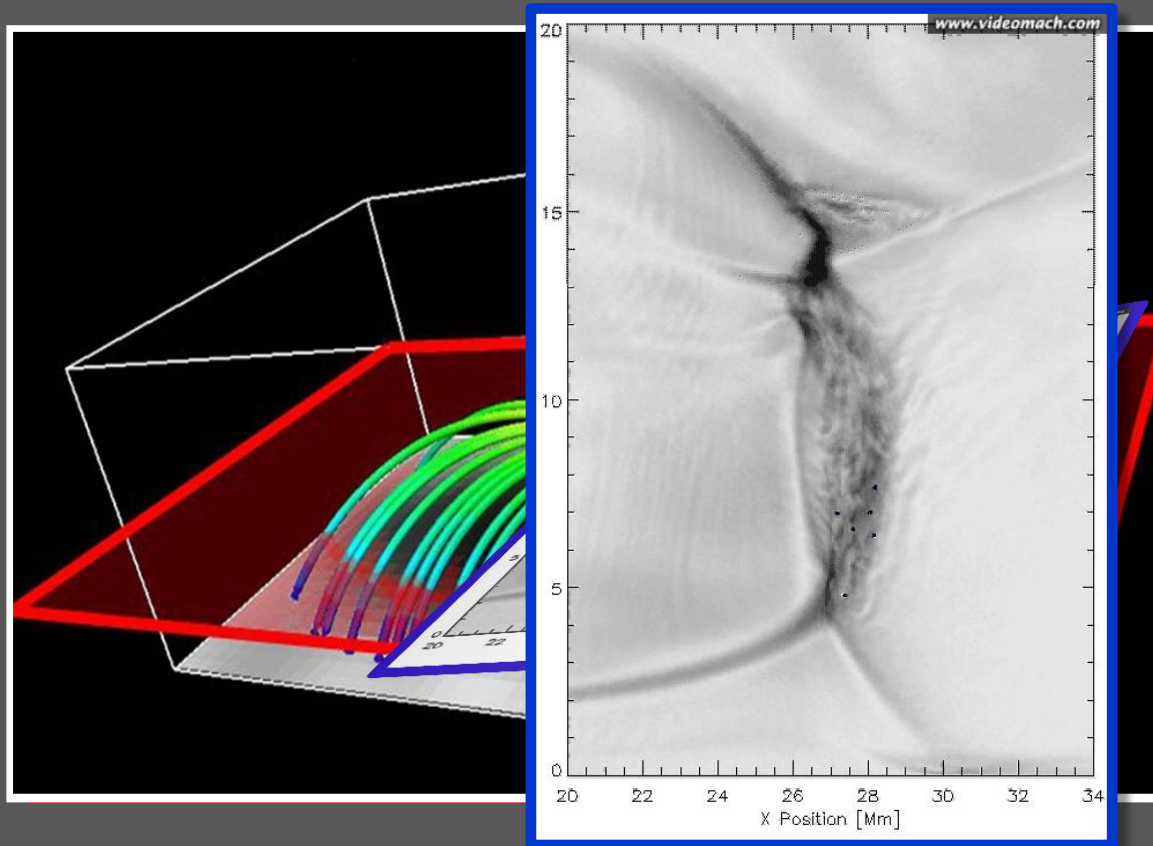






# Record Breaking: 135 billion particles

- ▣ Solar cut-out, potential field extrapolation, MHD model, particle-model-cutout, particle acceleration



# PRACE Specifics

- ▣ Intermittent Tier-0 access typical
  - 1 year projects (until now)
    - ▣ multi-year (2 yr!) access from 5th call
  - Needs to be complemented by Tier-1 & local facilities
  - Development, training, ...
  
- ▣ Front line research generates TB – PB of data
  - Needs to be **analyzed**, moved & **archived**
  - **Access** should be made possible for peers
    - ▣ Sub-projects, derived data ...



# Analysis Requirements

## ▣ Analysis

- Huge RAM size
- Multi-core
- High I/O bandwidth

⇒ Virtual Shared Memory (ScaleMP) is ideal

## ▣ At DCSC-KU

- ScaleMP system, similar to 'Gordon' at San Diego (SDSC)
- So far: 4 nodes = 48 cores, 256 GB ram, 200 TB disks
- Expandable ....

# Archival Requirement

- ▣ PRACE related
  - Data to be moved off PRACE system 2 mo after project end
  
- ▣ Simulation pipe-line – *must* have an open end!
  - Runs (PRACE, NASA/Ames, DCSC-KU)
  - Analysis (local – at FZ-J, Ames, DCSC-KU)
  - Restarts, continuations, spin-off, ...
  - Peer access – by collaboration only, so far
  - Archival – Hierarchical Storage Management preferred
    - ▣ Absolutely necessary to sustain the pipe line
    - ▣ Unfortunately not yet implemented at DCSC-KU

# Peer Access

- ▣ Collaborations (or not), sub-projects, spin-off, ...
  - Should / must allow peers to login, use *analysis* facilities
  - This avoids massive waste of bandwidth & human time
- ▣ Easy to combine with existing analysis facilities!

# Summary

- ▣ PRACE specifics
  - Intermittent (1-2 yr) Tier-0 access typical
    - ▣ Needs to be complemented by Tier-1 & local resources
      - Development, training, ...
  - Front-line science  $\Rightarrow$  TB – PB of data generated
    - ▣ Needs to be moved & archived
- ▣ PRACE-related requirements
  - Local (national?) data facilities
    - ▣ **Analysis** (virtual shared memory preferred)
    - ▣ **Archival** (Hierarchical Storage Management preferred)
    - ▣ **Access**: external **analysis access by peers**